JOT JOURNAL OF TURBULENCE

# Evaluation of turbulence models in MHD channel and square duct flows

Journal:	Journal of Turbulence
Manuscript ID:	JOT-2010-0050
Manuscript Type:	Original Manuscript
Date Submitted by the Author:	22-Jun-2010
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Keywords:	Electro- and Magneto-Hydrodynamic Turbulence, Direct Numerical Simulation, Turbulence Modeling: Reynolds averaged, Reynolds Averaged Navier-Stokes



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# ABSTRACT

In the present paper, several turbulence models have been evaluated in a channel and a square duct flow with and without a magnetic field by comparing the predictions with Direct Numerical Simulations (DNS) data. The various turbulence models include high and low Reynolds number versions of the k- $\varepsilon$  model, low and high-Re versions of the Reynolds-stress transport models with and without modifications for the application of a magnetic field. The simulations are performed using the FLUENT computer program. The additional source terms for the magnetic effects on turbulence have been included through user-defined functions. A systematic assessment of the predicted mean flow, turbulence quantities, frictional losses, and computational costs of the various turbulence models is presented.

All the models predict mean axial velocity reasonably well, but the predictions of turbulence parameters are less accurate. Velocity predictions are worse for the square duct flow due to secondary flows generated by the turbulence. The implementation of the MHD sources generally improves predictions in MHD flows, especially for low-Re k- $\epsilon$  models. The high-Re models using the wall treatments show little improvement, perhaps due to the lack of MHD effects in the wall formulations. Finally, at low Reynolds numbers, the Lam-Bremhorst (LB) low-Re k- $\epsilon$  model was found to give better predictions than other models for both hydrodynamic and magnetic field influenced turbulent flows.

Keywords: Turbulence, DNS, RANS, MHD, magnetic field

# **1. INTRODUCTION**

Reynolds-averaged Navier-Stokes (RANS) simulations are widely used to optimize various industrial flows because of their low computational cost. However, it is well-known that their accuracy in complex flows is limited by the difficulties in modeling the complex turbulence interactions through transport equations for the mean flow variables [1]. Significant effort has already been devoted to validation, improvement, and custom tailoring of these models of turbulent flows for different classes of flows [2-8]. This is usually done through comparisons with experimental data. However, with the availability of Direct Numerical Simulation (DNS)

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Despite the importance of magnetic fields in material processing, very limited work [12-15] exists on improving and testing turbulence models to include the effects of a magnetic field on the turbulence. A few modified models with magnetic field effects have been tested in channel flow/rectangular duct flow with a partial magnetic field (low-Re k- $\varepsilon$  and RSM) [12-13], pipe flow (low-Re k- $\varepsilon$ ) [14] and free surface channel flow (k- $\varepsilon$ ) [15]. The modifications proposed in the latter two of these studies (pipe flow [14] and free surface channel flow [15]) were based upon bulk properties of the flow and cannot be generalized to other flows. The first two studies (k- $\varepsilon$  and RSM, [12-13]) relate the magnetic field generated source terms in the turbulent transport equations to the local properties, and therefore can be generalized to other flows. However, these models have been so far tested only in a turbulent channel flow and in a rectangular duct with a partial magnetic field. For the rectangular duct with a partial magnetic field only the mean velocity was compared. The mean velocity obtained with this model was reported to show better agreement with measurements but no comparisons are available for turbulence quantities [12].

The present work reports a systematic assessment of a number of turbulence models, and their variants, for MHD flow in two representative geometries: a) channel flow, and b) a square duct flow. Confined internal flows through long pipes and ducts are relevant in many commercial flows. The square duct flow is more complicated to predict because of the turbulence-driven secondary flows [16]. The various models considered are: a) 3 variants of high-Re two-equation models (Standard k- $\epsilon$  (SKE) [17], RNG k- $\epsilon$  (RNG) [18], Realizable k- $\epsilon$  (RKE) [19], b) 6 low-Re k- $\epsilon$  models (Abid [20], Lam-Bremhorst (LB) [21], Launder-Sharma (LS) [22], Yang-Shih (YS) [23], Abe-Kondoh-Nagano (AKN) [24], and Chang-Hsieh-Chen (CHC) [25-26]) and c) 2 second-momentum closure Reynolds Stress Models with Linear Pressure Strain (RSM-LPS) and Stress-Omega (RSM-S $\omega$ ) [27-31]) models along with standard wall functions [32], non-equilibrium wall functions [33], and two-layer wall treatment combined with single-blended wall function (enhanced wall treatment) [34-35, 30]. The simulations have been performed using FLUENT [30] and the effect of magnetic field on turbulence, as given by Kenjereš and Hanjalić

[12-13], has been incorporated through additional source terms using User-Defined Functions (UDF). Mean velocities, turbulent kinetic energy (TKE), RMS of velocity fluctuations, MHD sources/sinks and frictional losses are compared against available DNS data in these two geometries.

# **2. TURBULENCE MODELS TESTED**

#### 2.1. Base Models

In the RANS approach, the ensemble averaged Navier-Stokes equations are written as [36-37]:

$$\frac{\partial \overline{u}_i}{\partial t} + \frac{\partial \overline{u}_i \overline{u}_j}{\partial x_j} = -\frac{1}{\rho} \frac{\partial \overline{p}}{\partial x_i} + \frac{\partial}{\partial x_j} \left( \nu \frac{\partial \overline{u}_i}{\partial x_j} \right) + \frac{\partial R_{ij}}{\partial x_j} + \overline{F}_L$$
(1)

where,  $R_{ij} = -\overline{u'_i u'_j}$ : Reynolds Stresses, and  $\overline{F}_L$  is the average Lorentz force due to magnetic field. Six of the nine components of the Reynolds stresses are independent.

Models using Boussinesq hypothesis relate the Reynolds stresses to the mean velocity gradients

and an isotropic eddy viscosity (i.e. 
$$R_{ij} = -\overline{u'_i u'_j} = v_t \left(\frac{\partial \overline{u}_i}{\partial x_j} + \frac{\partial \overline{u}_j}{\partial x_i}\right) - \frac{2}{3}k\delta_{ij}$$
) (SKE, RNG, RKE, and

low-Re k- $\varepsilon$  models etc.). On the other hand, the Reynolds-stress transport models (RSM) determine these stresses by solving six more transport equations for the six Reynolds stresses, so are more computationally expensive. The various models tested in this study are first given below without changes for magnetic field effects. The modifications for the presence of a magnetic field are subsequently described.

## 2.1.1. Standard k-ε model (SKE)

In this two equation model [17], the transport equations for turbulent kinetic energy (k) and its dissipation ( $\epsilon$ ) are written as,

$$\frac{\partial \rho k}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k k \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_k} \right) \frac{\partial k}{\partial x_j} \right] + G_k - \rho \varepsilon$$
(2)

$$\frac{\partial \rho \varepsilon}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k \varepsilon \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_\varepsilon} \right) \frac{\partial \varepsilon}{\partial x_j} \right] + C_{\varepsilon 1} \frac{\varepsilon}{k} G_k - \rho C_{\varepsilon 2} \frac{\varepsilon^2}{k}$$
(3)

Where, 
$$k = \overline{u_i'u_i'}/2$$
,  $\varepsilon = v \frac{\overline{\partial u_i'}}{\partial x_j} \frac{\partial u_i'}{\partial x_j}$ , and  $G_k = -\rho \overline{u_i'u_j'} \frac{\partial \overline{u_j}}{\partial x_i}$   
 $C_\mu = 0.09$ ,  $C_{\varepsilon 1} = 1.44$ ,  $C_{\varepsilon 2} = 1.92$ ,  $\sigma_k = 1.0$ ,  $\sigma_{\varepsilon} = 1.3$ ,  $v_t = C_\mu \frac{k^2}{\varepsilon}$ .

In  $G_k$ ,  $\overline{u'_i u'_j}$  is closed by the eddy viscosity model.

# 2.1.2. RNG k-ɛ model (RNG)

The RNG k- $\varepsilon$  turbulence model [18] includes an additional term in the  $\varepsilon$  equation, and uses different turbulent Prandtl numbers in the k and  $\varepsilon$  equations, as follows.

$$\frac{\partial \rho k}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k k \right) = \frac{\partial}{\partial x_j} \left[ \alpha_k \mu_{eff} \, \frac{\partial k}{\partial x_j} \right] + G_k - \rho \varepsilon \tag{4}$$

$$\frac{\partial \rho \varepsilon}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k \varepsilon \right) = \frac{\partial}{\partial x_j} \left[ \alpha_{\varepsilon} \mu_{eff} \, \frac{\partial \varepsilon}{\partial x_j} \right] + C_{\varepsilon 1} \frac{\varepsilon}{k} G_k - \rho C_{\varepsilon 2} \, \frac{\varepsilon^2}{k} - R_{\varepsilon} \tag{5}$$

With  $C_{\varepsilon 1} = 1.42$ ,  $C_{\varepsilon 2} = 1.68$ , and  $R_{\varepsilon}$  is the additional term in the  $\varepsilon$  equation for rapidly strained flows. The inverse Prandtl numbers ( $\alpha_k$  and  $\alpha_{\varepsilon}$ ) are calculated using the following formula derived from RNG theory,

$$\left|\frac{\alpha - 1.3929}{\alpha_o - 1.3929}\right|^{0.6321} \left|\frac{\alpha + 2.3929}{\alpha_o + 2.3929}\right|^{0.3679} = \frac{\mu}{\mu_{eff}}, \ \alpha_o = 1$$
(6)

Based upon RNG theory, a differential formulation for effective viscosity for low-Re effects is defined as;

$$d\left(\frac{\rho^2 k}{\sqrt{\varepsilon\mu}}\right) = 1.72 \frac{\hat{\nu}}{\sqrt{\hat{\nu}^3 - 1 + C_{\nu}}} d\hat{\nu}$$
(7)

Where,  $\hat{\nu} = \frac{\mu_{eff}}{\mu}$  and  $C_{\nu} = 100$ 

This equation can be integrated for  $\hat{v}$  and the integration constant can be calculated under the condition that  $\hat{v} = 1$ , when k = 0. In high-Re limit, this formulation reduces to the same expression as in SKE (i.e.  $v_t = C_{\mu} \frac{k^2}{\varepsilon}$ ,  $C_{\mu} = 0.09$ ).

The additional term  $(R_{\varepsilon})$  in  $\varepsilon$  equation accounting for the effects of rapid strain is defined as,

$$R_{\varepsilon} = \frac{C_{\mu}\rho\eta^{3}\left(1-\eta/\eta_{0}\right)}{1+\beta\eta^{3}}\frac{\varepsilon^{2}}{k}, \ \eta = \frac{Sk}{\varepsilon}, \eta_{0} = 4.38, \beta = 0.012$$
(8)

where, 
$$S = \sqrt{2s_{i,j}s_{i,j}}$$
,  $s_{i,j} = \frac{1}{2} \left( \frac{\partial \overline{u}_i}{\partial x_j} + \frac{\partial \overline{u}_j}{\partial x_i} \right)$ 

# 2.1.3. Realizable k-ε model (RKE)

The realizable k- $\varepsilon$  (RKE) model has been proposed by Shih et al [19]. This model has a realizable formulation for Reynolds normal stresses (i.e. positivity) and does not violate Schwarz inequality  $(\overline{u'_a u'_b}^2 \le \overline{u'_a^2 u'_b})$  in highly strained flows. The realizable formulation of Reynolds stresses is obtained by sensitizing the constant  $(C_{\mu})$  of eddy viscosity equation to the mean flow, k and  $\varepsilon$ . In addition to variable  $C_{\mu}$ , RKE model also has a new formulation for dissipation rate ( $\varepsilon$ ) derived from the exact mean-square vorticity fluctuation equation (because  $\varepsilon = v \overline{\omega_i \omega_i}$ ,  $\omega_i = \frac{\partial u'_k}{\partial x_j} - \frac{\partial u'_j}{\partial x_k}$ ). The governing equations for k- and  $\varepsilon$ - in RKE model are .

given as,

$$\frac{\partial \rho k}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k k \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_k} \right) \frac{\partial k}{\partial x_j} \right] + G_k - \rho \varepsilon \tag{9}$$

$$\frac{\partial \rho \varepsilon}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k \varepsilon \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_\varepsilon} \right) \frac{\partial \varepsilon}{\partial x_j} \right] + \rho C_1 S \varepsilon - \rho C_2 \frac{\varepsilon^2}{k + \sqrt{\nu \varepsilon}}$$
(10)

$$C_2 = 1.9$$
,  $\sigma_k = 1.0$ ,  $\sigma_{\varepsilon} = 1.2$ ,  $C_1 = \max\left(0.43, \frac{\eta}{\eta+5}\right)$ , where  $\eta$  is as defined in Eq-8 for the

RNG model.  $G_k$  is defined the same way as in SKE and RNG. The eddy viscosity is also defined

same as before, i.e. 
$$\mu_t = \rho C_{\mu} \frac{k^2}{\varepsilon}$$
.

For RKE, 
$$C_{\mu} = \frac{1}{A_0 + A_s \frac{kU^*}{\varepsilon}}$$
, where  $U^* = \sqrt{s_{i,j}s_{i,j} + \tilde{\Omega}_{i,j}\tilde{\Omega}_{i,j}}$  (11)

$$\tilde{\Omega}_{i,j} = \Omega_{i,j} - 2\varepsilon_{ijk}\omega_k , \ \Omega_{i,j} = \overline{\Omega}_{i,j} - \varepsilon_{ijk}\omega_k , \ \overline{\Omega}_{i,j} = \frac{1}{2} \left( \frac{\partial \overline{u}_i}{\partial x_j} - \frac{\partial \overline{u}_j}{\partial x_i} \right)$$

Where,  $\varepsilon_{ijk}$  is Levi-Civita symbol;  $\varepsilon_{ijk} = \begin{cases} 1 & \text{if}(i, j, k) \text{ are cyclic} \\ -1 & \text{if}(i, j, k) \text{ are anticyclic} \\ 0 & \text{otherwise} \end{cases}$ 

The constants,  $A_0 = 4.04$  and  $A_s = \sqrt{6}\cos(\phi)$ ,  $\phi = \frac{1}{3}\cos^{-1}(\sqrt{6}W)$ ,  $W = \frac{s_{i,j}s_{j,k}s_{k,i}}{\tilde{S}^3}$ ,  $\tilde{S} = \sqrt{s_{i,j}s_{i,j}}$ 

# 2.1.4. Low-Re k-ε models

Several low-Re k- $\varepsilon$  models [20-26] have been proposed (Abid [20], Launder-Sharma (LS) [21], Lam-Bremhorst (LB) [22], Yang-Shih (YS) [23], Abe-Kondoh-Nagano (AKN) [24], and Chang-Hsieh-Chen (CHC) [25-26]). These models use damping functions to make the model valid in the near wall regions. The general k and  $\varepsilon$  equations for low-Re k- $\varepsilon$  models can be written as:

$$\frac{\partial \rho k}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k k \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_k} \right) \frac{\partial k}{\partial x_j} \right] + G_k - \rho \varepsilon - \rho D$$
(12)

$$\frac{\partial \rho \varepsilon}{\partial t} + \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k \varepsilon \right) = \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_t}{\sigma_\varepsilon} \right) \frac{\partial \varepsilon}{\partial x_j} \right] + f_1 C_1 \frac{\varepsilon}{k} G_k - \rho f_2 C_2 \frac{\varepsilon^2}{k} + \rho E$$
(13)

$$\mu_{t} = \rho f_{\mu} C_{\mu} \frac{k^{2}}{\varepsilon}$$
(14)

Damping functions, wall boundary conditions and various constant for different low-Re k-ε models are given in Table 1 and Table 2.

# 2.1.5. Reynolds Stress Model (RSM)

The exact transport equation for the six independent Reynolds stresses  $(\overline{u'_iu'_j})$  in RSM can be written as [27-31]:

$$\frac{\partial}{\partial t} \left( \rho \overline{u'_{i} u'_{j}} \right) + \frac{\partial}{\partial x_{k}} \left( \rho \overline{u}_{k} \overline{u'_{i} u'_{j}} \right) = P_{ij} + D_{ij}^{L} + D_{ij}^{T} + \phi_{ij} - \varepsilon_{ij}$$
(15)

$$P_{ij} = -\rho \left( \overline{u_i' u_k'} \frac{\partial \overline{u}_j}{\partial x_k} + \overline{u_j' u_k'} \frac{\partial \overline{u}_i}{\partial x_k} \right) \text{(I: Production), } D_{ij}^L = \frac{\partial}{\partial x_k} \left( \mu \frac{\partial}{\partial x_k} \left( \overline{u_i' u_j'} \right) \right) \text{(II: Molecular diffusion),}$$

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$$D_{ij}^{T} = -\frac{\partial}{\partial x_{k}} \left( \rho \overline{u_{i}' u_{j}' u_{k}'} + \overline{p\left(\delta_{kj} u_{i}' + \delta_{ik} u_{j}'\right)} \right) \quad \text{(III: Turbulent diffusion),} \quad \phi_{ij} = p \left( \frac{\partial u_{i}'}{\partial x_{j}} + \frac{\partial u_{j}'}{\partial x_{i}} \right) \quad \text{(IV)}$$

Pressure strain),  $\varepsilon_{ij} = 2\mu \frac{\partial u'_i}{\partial x_k} \frac{\partial u'_j}{\partial x_k}$  (V: Dissipation), where  $\delta_{ij} = 1$ , if i=j, else 0.

Of these five terms, the last three  $(D_{ij}^T, \phi_{ij} \text{ and } \varepsilon_{ij})$  require modeling, with the pressure strain  $(\phi_{ij})$  and dissipation  $(\varepsilon_{ij})$  considered to be critical [28].

The turbulent diffusion term (i.e.  $D_{ij}^T$ , III) is modeled the same way as the molecular diffusion term (Lien and Leschziner [38]):

$$D_{ij}^{T} = \frac{\partial}{\partial x_{k}} \left( \frac{\mu_{t}}{\sigma_{k}} \frac{\partial}{\partial x_{k}} \left( \overline{u_{i}' u_{j}'} \right) \right)$$
(16)

where  $\mu_t = \rho C_\mu \frac{k^2}{\varepsilon}$ ,  $C_\mu = 0.09$ ,  $\sigma_k = 0.82$ 

To model dissipation tensor term (i.e.  $\varepsilon_{ij}$ ), a transport equation similar to standard k- $\varepsilon$  model (with  $\sigma_{\varepsilon} = 1.0$ ) for dissipation rate (i.e.  $\varepsilon$ ) is solved. The dissipation tensor is defined from dissipation rate as:

$$\varepsilon_{ij} = \frac{2}{3} \delta_{ij} \rho \varepsilon \tag{17}$$

The main difference in different RSM models is due to the handling of pressure strain term ( $\phi_{ij}$ ). Gibson and Launder [27], Launder [28, 40], Fu et al. [39], Launder and Shima [41] and Wilcox [31] proposed different ways to model this term in high- and low-Re versions of RSM model. In the current work, low- and high-Re versions of Linear Pressure Strain model and low-Re stress omega model formulations to handle pressure strain term have been used [30]. The high-Reynolds number version of Linear Pressure Strain formulation is used in conjunction with the standard or non-equilibrium wall functions. When the enhanced wall treatment is used, the modified Linear Pressure Strain model formulation incorporating the low-Re effects is used. These are briefly described below,

In the linear pressure strain model (RSM-LPS), the pressure strain term is decomposed into three components,

$$\phi_{ij} = \phi_{ij1} + \phi_{ij2} + \phi_{ijw} \tag{18}$$

$$\phi_{ij1} = -C_1 \rho \frac{\varepsilon}{k} \left( \overline{u'_i u'_j} - \frac{2}{3} \delta_{ij} k \right), \ C_1 = 1.8$$
(19)

$$\phi_{ij2} = -C_2 \left( \left( P_{ij} - C_{ij} \right) - \frac{2}{3} \delta_{ij} \left( \frac{P_{kk}}{2} - \frac{C_{kk}}{2} \right) \right), \tag{20}$$

where,  $C_2 = 0.6$  and  $C_{ij} = \frac{\partial}{\partial x_k} \left( \rho \,\overline{u}_k \,\overline{u'_i u'_j} \right)$  (convection term)

$$\phi_{ijw} = C_1' \frac{\varepsilon}{k} \left( \overline{u'_k u'_m} n_k n_m \delta_{ij} - \frac{3}{2} \overline{u'_i u'_k} n_j n_k - \frac{3}{2} \overline{u'_j u'_k} n_i n_k \right) \frac{C_l k^{3/2}}{\varepsilon d}$$

$$+ C_2' \left( \phi_{km2} n_k n_m \delta_{ij} - \frac{3}{2} \phi_{ik2} n_j n_k - \frac{3}{2} \phi_{jk2} n_i n_k \right) \frac{C_l k^{3/2}}{\varepsilon d}$$

$$(21)$$

 $C_1 = 0.5$ ,  $C_2 = 0.3$ ,  $C_l = C_{\mu}^{3/4} / \kappa$ ,  $C_{\mu} = 0.09$ ,  $\kappa = 0.42$ , *d* is the normal distance to the wall.  $n_k$  is the  $x_k$  component of unit normal vector.

In low-Re version of RSM-LPS (which is used with enhanced wall treatment), the constants ( $C_1$ ,  $C_2$ ,  $C_1$  and  $C_2$ ) are sensitized to Reynolds stress invariants and turbulent Reynolds number ( $\operatorname{Re}_T = \frac{\rho k^2}{\mu \varepsilon}$ ) [41].  $C_1 = 1 + 2.58A\sqrt{A_2} \left(1 - \exp\left(-\left(0.0067 \operatorname{Re}_T\right)^2\right)\right), C_2 = 0.75\sqrt{A}, C_1 = -\frac{2}{3}C_1 + 1.67,$   $C_2 = \max\left(\frac{\frac{2}{3}C_2 - \frac{1}{6}}{C_2}, 0\right)$ Where,  $A = 1 - \frac{9}{8}(A_2 - A_3), A_2 = a_{ik}a_{ki}, A_3 = a_{ik}a_{kj}a_{ji}, a_{ij} = -\left(\frac{-\rho u_i' u_j' + \frac{2}{3}\rho k\delta_{ij}}{\rho k}\right)$  Besides RSM with linear pressure strain formulation, in one calculation of a low-Re non-MHD channel flow, the RSM with low-Re stress omega (RSM-S $\omega$ ) formulation given by Wilcox [31] has also been used. Details of this model can be found in [30] and [31].

## 2.2. Near-Wall treatment

Near-wall treatment is very important in wall-bounded turbulent flows. Walls have high velocity gradients and thus are the main source of turbulence production. Near the wall there are three main regions [36]: i) Viscous sublayer where molecular viscosity is dominant, ii) Buffer region where molecular and turbulence effects are important and overlap together and iii) Fully turbulent region (log-law region).

These wall regions are differently handled in different models. The low-Re models (i.e. Abid, LB, LS, YS, AKN, CHC, RSM-S $\omega$  with low Re-correction) use damping functions and need a fine grid to integrate up to viscous sublayer ( $y^+ = \frac{yu_r}{v}(\tau_w = \rho u_r^2) <=1$ ) [42]. In high-Re models (i.e. RKE, SKE, RNG, RSM etc.), the near-wall region is usually handled in two ways [30-33]: i) wall function approach without resolving the buffer and the viscous sublayers (applicable for  $30 < y^+ = \frac{yu_r}{v}(\tau_w = \rho u_r^2) <500$ : Standard Wall Function (SWF) and Non-Equilibrium Wall Functions (NEWF)), ii) Two-layer model for  $\varepsilon$  and turbulent viscosity with single blended law of wall for mean velocity (Enhanced Wall Treatment (EWT)). Formulations for the different wall treatment methods (SWF, NEWF and EWT) are given below.

#### 2.2.1. Standard wall function (SWF)

Launder and Spalding [32, 30] gave the standard law of wall for mean velocity as;

$$U^* = \frac{1}{\kappa} \ln(Ey^*), \ \kappa = 0.418, \ E = 9.79, \ C_{\mu} = 0.09$$
(22)

Where,  $U^* = \frac{U_p C_{\mu}^{1/4} k_p^{1/2}}{\tau_w / \rho}$ ,  $y^* = \frac{\rho C_{\mu}^{1/4} k_p^{1/2} y_p}{\mu}$ , (Note: in equilibrium boundary layer  $y^*$  and

 $y^+ = \frac{yu_{\tau}}{v}$  are approximately equal) (subscript *p* stands for the cell center next to wall).  $U_p$  and

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 $k_p$ , and  $y_p$  are the TKE, tangential velocity and distanced of cell center from wall in the cell next to the wall respectively.  $\tau_w$  is the wall shear stress.

At the wall, the normal derivative of TKE is taken zero (i.e.  $\frac{\partial k}{\partial n} = 0$ ) and assuming rate of TKE production equal to rate of dissipation, the value of dissipation in the cell next to the wall can be calculated as,

$$\varepsilon_p = \frac{C_{\mu}^{3/4} k_p^{3/2}}{\kappa y_p}$$

# 2.2.2. Non-equilibrium wall function (NEWF)

Kim and Choudhury [33, 30] sensitized the log-law mean velocity of SWF with pressure and proposed a two layer NEWF approach for production and dissipation of turbulence. In this formulation,

$$\frac{\tilde{U}C_{\mu}^{1/4}k^{1/2}}{\tau_{w}/\rho} = \frac{1}{\kappa} \ln\left(E\frac{\rho C_{\mu}^{1/4}k^{1/2}y}{\mu}\right)$$
(23)

$$\tilde{U} = U - \frac{1}{2} \frac{\partial p}{\partial x} \left( \frac{y_{\nu}}{\rho \kappa \sqrt{k}} \ln \left( \frac{y}{y_{\nu}} \right) + \frac{y - y_{\nu}}{\rho \kappa \sqrt{k}} + \frac{y_{\nu}^{2}}{\mu} \right), \quad y_{\nu} = \frac{\mu y_{\nu}^{*}}{\rho C_{\mu}^{1/4} k_{p}^{1/2}}, \quad y_{\nu}^{*} = 11.225.$$

 $y_{\nu}$  is viscous sublayer thickness. Now, a two layer concept is used in the cell next to wall to

calculate k and 
$$\varepsilon$$
.  $\tau_t = \begin{cases} 0, \ y < y_v \\ \tau_w, \ y > y_v \end{cases}$ ,  $k = \begin{cases} \left(\frac{y}{y_v}\right)^2 k_p, \ y < y_v \\ k_p, \ y > y_v \end{cases}$ ,  $\varepsilon = \begin{cases} \frac{2\nu k}{y^2}, \ y < y_v \\ \frac{k^{3/2}}{C_l^* y}, \ y > y_v \end{cases}$ ,  $C_l^* = \kappa C_{\mu}^{-3/4}$ 

Using above profiles, the cell-average production ( $G_k$ ) of turbulent kinetic energy and dissipation rate ( $\varepsilon$ ) can be calculated as:

$$G_k = \frac{1}{y_n} \int_0^{y_n} \tau_t \frac{\partial U}{\partial y} dy = \frac{1}{\kappa y_n} \frac{\tau_w^2}{\rho C_\mu^{1/4} k_p^{1/2}} \ln\left(\frac{y_n}{y_v}\right), \text{ and } \varepsilon = \frac{1}{y_n} \int_0^{y_n} \varepsilon dy = \frac{1}{y_n} \left(\frac{2\nu}{y_v} + \frac{k_p^{1/2}}{C_l^*} \ln\left(\frac{y_n}{y_v}\right)\right) k_p$$

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With  $G_k$  and  $\varepsilon$ , the TKE equation is solved in the domain with  $\frac{\partial k}{\partial n} = 0$  at the wall.  $y_n$  is the wall normal height of the cell  $(y_n = 2y_p)$ 

# 2.2.3. Enhanced wall treatment (EWT)

The EWT uses a two layer approach for eddy viscosity and dissipation rate based upon turbulent Reynolds number ( $\operatorname{Re}_{y} = \frac{\rho \sqrt{k} y}{\mu}$ , y is the normal distance from cell center to the wall) [30]. In the viscous region (i.e.  $\operatorname{Re}_{y} < \operatorname{Re}_{y}^{*}$ ,  $\operatorname{Re}_{y}^{*} = 200$ ), a one-equation model of Wolfstein [35] and in turbulent region (i.e.  $\operatorname{Re}_{y} > \operatorname{Re}_{y}^{*}$ ) the selected turbulence model is used. In viscous region, the momentum equation and TKE equations are solved as usual but the eddy viscosity and dissipation rates are calculated as:

$$\mu_{t,2} = \rho C_{\mu} l_{\mu} \sqrt{k} , \ l_{\mu} = y C_l^* \left( 1 - \exp(-\operatorname{Re}_y / A_{\mu}) \right)$$
(24)  
$$\varepsilon = \frac{k^{3/2}}{l_{\varepsilon}}, \ l_{\varepsilon} = y C_l^* \left( 1 - \exp(-\operatorname{Re}_y / A_{\varepsilon}) \right), \text{ where } C_l^* = \kappa C_{\mu}^{-3/4}, \ A_{\mu} = 70, \ A_{\varepsilon} = 2C_l^*.$$

Further, the eddy viscosity of viscous region is blended with the fully turbulent viscosity to give smooth behavior in between viscous and fully turbulent regions as follows:

$$\mu_{t,enhanced} = \lambda_{\varepsilon} \mu_t + (1 - \lambda_{\varepsilon}) \mu_{t,2}$$
 Where,  $\mu_t = \rho C_{\mu} \frac{k^2}{\varepsilon}$ ,  $C_{\mu} = 0.09$ , same blending is performed

for  $\varepsilon$ . The blending function is defined as:  $\lambda_{\varepsilon} = \frac{1}{2} \left( 1 + \tanh\left(\frac{\operatorname{Re}_{y} - \operatorname{Re}_{y}^{*}}{A}\right) \right)$ , A is the width of the

blending function,  $A = \left| \frac{\Delta \text{Re}_y}{\tanh(0.98)} \right|$ ,  $\Delta \text{Re}_y$  is assigned a value in between 5% to 20% of  $\text{Re}_y^*$  to

give smooth behavior.

Furthermore, for the mean tangential velocity, a blended single wall law is used [30, 34]. The blended single wall law is defined as,

$$U^{+} = e^{\Gamma} U^{+}_{\text{laminar}} + e^{1/\Gamma} U^{+}_{\text{turbulent}} \text{ and } \frac{dU^{+}}{dy^{+}} = e^{\Gamma} \frac{dU^{+}_{\text{laminar}}}{dy^{+}} + e^{1/\Gamma} \frac{dU^{+}_{\text{turbulent}}}{dy^{+}}$$
(25)

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Where, the blending function is  $\Gamma = -\frac{a(y^+)^4}{1+by^+}$ , a = 0.01, and b = 5. For turbulent region, the

wall law as given by White and Cristoph [43] and Huang et al [44] with the effect of pressure is as follows:

$$\frac{dU_{\text{turbulent}}^{+}}{dy^{+}} = \frac{1}{\kappa y^{+}} \left(S'\right)^{1/2}, \ S' = \begin{cases} 1 + \alpha y^{+}, \ y^{+} < y_{s}^{+} \\ 1 + \alpha y_{s}^{+}, \ y^{+} > = y_{s}^{+} \end{cases}, \ y_{s}^{+} = 60, \text{ and } \alpha = \frac{\mu}{\rho^{2} u_{\tau}^{3}} \frac{\partial p}{\partial x}$$

The above ordinary differential equation can be integrated to find  $U_{turbulent}^+$ , especially in case of  $\alpha = 0$ , it will give the log-law of wall.

For laminar region,  $\frac{dU_{\text{laminar}}^+}{dy^+} = 1 + \alpha y^+$ , By integration  $U_{\text{laminar}}^+ = y^+ \left(1 + \frac{\alpha}{2} y^+\right)$ 

The TKE is solved in the whole domain with  $\frac{\partial k}{\partial n} = 0$  at the wall, and the  $G_k$  term in TKE equation is calculated using the velocity gradient (Eq-25) consistent with single wall law as given above.

# 2.2.4. Wall treatment in RSM model for Reynolds stresses

RSM model solves for individual Reynolds stresses and therefore it needs boundary conditions for Reynolds stresses in addition to the above wall treatment procedures. With SWF and NEWF, turbulent kinetic energy is calculated using  $k = \frac{\overline{u'_i u'_i}}{2}$  away from the wall but in the cells next to the walls, a transport equation, similar to SKE, for TKE (with  $\sigma_k = 0.82$ ) is solved with normal derivative of k equal to zero at the wall. Afterwards, the following equations are used to calculate individual Reynolds stresses in the cells next to the wall (derived based upon equilibrium of Reynolds stresses, i.e. production=dissipation) [30].

$$\frac{\overline{u_{i}'u_{i}'}}{k} = 1.098, \ \frac{\overline{u_{\eta}'u_{\eta}'}}{k} = 0.247, \ \frac{\overline{u_{\lambda}'u_{\lambda}'}}{k} = 0.655, \ -\frac{\overline{u_{i}'u_{\eta}'}}{k} = 0.255$$
(26)

Where, subscript t,  $\eta$  and  $\lambda$  stands for local tangential, normal and binormal coordinates respectively. With EWT, the normal derivatives of Reynolds stresses are taken zero at the wall.

# 2.3. MHD formulations

The movement of a conducting fluid through an applied magnetic field induces a current, which generates a Lorentz force [45] that tends to oppose the flow. Two different modeling approaches are used to model fluid flow with MHD, depending on the importance of coupling between the applied and induced magnetic fields.

When the Magnetic Reynolds number,  $\operatorname{Re}_{m} = vL(\mu\sigma)$ , is <1 (such as for liquid metals), the induced magnetic field is negligible relative to the applied field, so the "electric potential method" is most efficient. Based on Ohm's law and conservation of charge, coupled equations for electric potential,  $\phi$ , and Lorentz force,  $\overline{F}_{L}$  can be solved as follows [45, 30].

$$\nabla^2 \phi = \nabla \cdot \left( \vec{v} \times \vec{B}_0 \right) \text{ and } \vec{F}_L = \sigma \left( -\nabla \phi + \vec{v} \times \vec{B}_0 \right) \times \vec{B}_0$$
(27)

In time varying fields, and when the induced current is significant, (i.e.  $\text{Re}_m > 1$ ), the "magnetic induction" method is best. Maxwell's equations are combined with Ohm's law to obtain a transport equation for the induced magnetic field,  $\vec{b}$  in terms of the total field,  $\vec{B}$  and the current density,  $\vec{J}$  [45, 30].

$$\frac{\partial \vec{b}}{\partial t} + (\vec{v} \cdot \nabla) \vec{b} = \frac{1}{\mu \sigma} \nabla^2 \vec{b} + ((\vec{B}_o + \vec{b}) \cdot \nabla) \vec{v} - (\vec{v} \cdot \nabla) \vec{B}_o - \frac{\partial \vec{B}_0}{\partial t} + \frac{1}{\mu \sigma} \nabla^2 \vec{B}_0$$

$$\vec{B} = \vec{B}_o + \vec{b} ; \qquad \vec{J} = \nabla \times \vec{B} / \mu ; \qquad \vec{F}_L = \vec{J} \times \vec{B}$$
(28)

In both methods, the Lorentz force is applied as a source term in the momentum equations.

# 2.4. Effect of magnetic field on turbulence in RANS turbulence models

Kenjereš and Hanjalić [12], Kenjereš, Hanjalić and Bal [13], Ji & Gardner [14], Smolentsev et al [15], and Galprin [46] improved the conventional non-MHD RANS turbulence models for the effect of the magnetic field on the turbulence in low magnetic Reynolds number liquid metal MHD flows. Ji & Gardner [14] proposed additional source terms for k- $\varepsilon$  turbulence model to account for the effect of magnetic field damping of turbulence. This extended k- $\varepsilon$  model was tested on the turbulent flow of an electrically conducting liquid in an insulated pipe. Velocity profiles, skin friction, temperature profiles, Nusselt numbers showed agreement with available experimental data for the range of Reynolds and Hartmann numbers. The biggest shortcoming of this model was the usage of bulk Stuart number (or interaction parameter, Ha<sup>2</sup>/Re) to define the turbulence damping terms making it a bulk flow dependent model and only applicable in the

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standard problems where bulk Stuart number can be easily defined. Smolentsev et al [15] proposed different source terms for k- $\epsilon$  models but again based up on the bulk flow Stuart number. The model was found to match experiments closely in free surface channel flow.

Galperin [46] was the first to propose a second-moment closure model for MHD turbulence, although this model was not numerically tested on conventional flows. Kenjereš and Hanjalić [12-13] proposed new source terms for k- $\varepsilon$  and second-moment closure models (RSM). The improved k- $\varepsilon$  model was validated with the DNS results in a channel flow under transverse magnetic field. After validation, the model was used in a 3-d developing rectangular duct flow with partial magnetic field and model was found performing well for mean velocities. No assessment for turbulence parameters was made in rectangular duct flow. Kenjereš and Hanjalić [12-13] also proposed a similar closure for  $\overline{u_i u_j}$  equations for MHD effects in RSM as proposed by Galperin [46]. This closure for RSM showed considerable improvement of results in a channel flow. The current study includes the models proposed by Kenjereš and Hanjalić's [12-13] for the channel and square duct flows. The following modifications were made to the models.

# 2.4.1. k-ε model :

k-equation: 
$$S_k^M = -\sigma B_0^2 k \exp\left(-C_1^M \frac{\sigma}{\rho} B_0^2 \frac{k}{\varepsilon}\right)$$
 (29)

$$\varepsilon\text{-equation:} \qquad S_{\varepsilon}^{M} = -\sigma B_{0}^{2} \varepsilon \exp\left(-C_{1}^{M} \frac{\sigma}{\rho} B_{0}^{2} \frac{k}{\varepsilon}\right)$$
(30)

Kenjereš and Hanjalić [12] used  $C_1^M = 0.025$  in channel flow under uniform transverse magnetic field. Same value of the constant ( $C_1^M$ ) has been used in the current work.

# 2.4.2. Reynolds Stress Model (RSM) MHD source terms:

After simplification for y-directional (vertical) magnetic field and some algebra the six independent Reynolds stress transport equations can be derived with the following MHD source terms;

$$\overline{w'w'} \text{-equation: } S_{w'w'}^{M} = \sigma \left( -2B_{y0} \overline{w'} \frac{\partial \phi'}{\partial x} - 2B_{y0}^{2} \overline{w'w'} \right)$$
(31)

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$$v'v'-\text{equation: } S^M_{v'v'} = 0 \tag{32}$$

$$\overline{u'u'} \text{-equation: } S_{u'u'}^{M} = \sigma \left( 2B_{y0}u' \frac{\partial \phi'}{\partial z} - 2B_{y0}^{2}\overline{u'u'} \right)$$
(33)

$$\overline{u'v'}\text{-equation: } S_{u'v'}^{M} = \sigma \left( B_{y0} \overline{v'} \frac{\partial \phi'}{\partial z} - B_{y0}^{2} \overline{u'v'} \right)$$
(34)

$$\overline{w'u'} \text{-equation: } S_{w'u'}^{M} = \sigma \left( -B_{y0} \overline{u' \frac{\partial \phi'}{\partial x}} + B_{y0} \overline{w' \frac{\partial \phi'}{\partial z}} - 2B_{y0}^{2} \overline{w'u'} \right)$$
(35)

$$\overline{w'v'} \text{-equation:} \quad S_{w'v'}^{M} = \sigma \left( -B_{y0} \overline{v'} \frac{\partial \phi'}{\partial x} - B_{y0}^{2} \overline{w'v'} \right)$$
(36)

Source term for scalar dissipation rate ( $\varepsilon$ ) is defined as [13];  $S_{\varepsilon}^{M} = \frac{1}{2} S_{ii}^{M} \frac{\varepsilon}{k}$  (37)

It can be seen that all the source terms due to the magnetic field are negatively correlated with the corresponding Reynolds stress therefore sinks to the Reynolds stresses. It is interesting to note that the magnetic field causes no direct sink to the Reynolds normal stress parallel to magnetic field (i.e.  $\overline{v'v'}$ ). The indirect suppression effect on  $\overline{v'v'}$  is via Reynolds shear stresses. In the above sinks, the terms involving correlation of velocity fluctuation with electric potential gradient require modeling and cannot be incorporated directly in RSM. Kovner and Levin [47] were the first to suggest a way to model electric potential-velocity correlation. Galperin [46] and later Kenjereš and Hanjalić [12-13] followed their method and came up with following formulation for the correlation;

$$\frac{\partial \phi'}{\partial x_k} = \beta \varepsilon_{kmn} u_m B_{n0} \Longrightarrow \overline{u_i' \frac{\partial \phi'}{\partial x_k}} = \beta \varepsilon_{kmn} \overline{u_i' u_m} B_{n0}$$
(38)

Where,  $\varepsilon_{kmn}$  is Levi-Civita symbol as defined in section in Realizable k- $\varepsilon$  model formulations. Galperin [46] proposed  $0 < \beta < 1$ . Kenjereš and Hanjalić [13] proposed  $\beta = 0.6$  via MHD channel flow. In the current work, the value of  $\beta$  as proposed by Kenjereš and Hanjalić is used.

The above discussed two formulations for k- $\varepsilon$  and RSM for the effect of magnetic field on turbulence have been implemented using a UDF with the magnetic induction and the electric potential methods [30]. More details on the various RANS type turbulence models (k- $\varepsilon$  models (Standard, RNG, Realizable) and RSM), various wall treatment approaches (standard and non-

equilibrium wall functions and enhanced wall treatment), magnetic induction and electric potential method for MHD calculations can be found in FLUENT manual [30].

# **3. DNS Data Bases**

Five DNS databases were used to assess the above models. The conditions for various DNS databases are given in Table 3.

#### 3.1. High-Reynolds Number Non-MHD channel flow

Satake et al [48] performed DNS calculations in a non-MHD channel at a bulk Reynolds number of ~45818 using 800 million nodes. The mean velocities, RMS of velocity fluctuations and turbulent kinetic energy budgets were reported. This non-MHD case was used as a base case to first evaluate the purely hydrodynamic models.

# 3.2. Low-Reynolds Number MHD and Non-MHD channel flows

The non-MHD channel flow data of Iwamoto et al [49] has been used to test performance of RANS models at lower Reynolds numbers. In his case,  $\text{Re}_{\tau} (=\delta u_{\tau}/v)=150$ , corresponding to bulk Re (= $2\delta W_b/v$ ,  $\delta$ : half channel height)=4586 was used. To test the models for MHD turbulence, the MHD channel case of Noguchi et al [50] ( $\text{Re}_{\tau} (=\delta u_{\tau}/v)=150$ , bulk Re (= $2\delta W_b/v$ )=4710, Ha (= $\text{sqrt}(\sigma/\rho v)B_0\delta$ )=6),  $\delta$ : half channel height) was used. Although these two cases have same applied mean streamwise pressure gradient corresponding to  $\text{Re}_{\tau}=150$  their bulk Reynolds numbers are different due to the differences in the frictional losses as a result of the magnetic field effects on the flow.

# 3.3. Low-Reynolds Number MHD and Non-MHD square duct flows

A GPU based code (CU-FLOW) [51] that has been previously used for DNS calculations in a non-MHD square duct has been extended for DNS calculations of a MHD square duct [52]. For the non-MHD case, ( $\text{Re}_{\tau}(=\text{Du}_{\tau}/v)=360$ , bulk Re ( $=\text{DW}_{b}/v$ )=5466), a duct of size of 1x1x8 non-dimensional units and 160x160x1024 control volumes (with 1% grid stretching in wall normal directions) were used. For the MHD case, ( $\text{Re}_{\tau}(=\text{Du}_{\tau}/v)=361$ , bulk Re ( $=\text{DW}_{b}/v$ )=5602, Ha ( $=\text{sqrt}(\sigma/\rho v)B_0D$ )=21.2)) a duct of size of 1x1x16 non-dimensional units with 128x128x512 control volumes (with 2% grid stretching in wall normal directions) were used. Both these

simulations were shown to give grid-independent solutions to the relevant equations. Note that the bulk Reynolds numbers were again different because of the magnetic field effects on the flow.

#### **4. COMPUTATIONAL DETAILS**

# 4.1. Computational Domain, Boundary Conditions and Numerical Method

Taking advantage of fully-developed flow with RANS models, the domain size was taken as 1x1x1 non-dimensional units for both the channel and the square duct. For the channel, the top and the bottom walls were electrically insulated with no-slip velocity conditions while the streamwise (z-) and spanwise (x-) directions were considered periodic. In the square duct, the four walls (top, bottom, right and left) were electrically insulated with no-slip velocity conditions whereas the streamwise direction (z-) is periodic. For the MHD calculations, the magnetic field was applied in the vertical (y-) direction. The simulations were carried out by fixing the bulk mean flow Reynolds number as given in Table 3 with the mean streamwise pressure gradient free to change. All the calculations were performed using FLUENT's steady-state segregated solver with SIMPLE algorithm for pressure-velocity coupling with either magnetic induction or electric potential methods for MHD calculations [30]. For each case, the results were ensured to be grid-independent by systematically increasing the number of control volumes until a grid-independent solution is obtained. All cases were converged such that the unscaled absolute residuals reached below  $10^{-3}$ .

# 4.2. Grids

For the high-Re calculations (case 1) with enhanced wall treatment, five grids with ten control volumes each in streamwise and spanwise directions were used. In the wall-normal direction, three uniform grids (consisting of 50, 80 and 130 control volumes) and two non-uniform grids (near-wall y+ = 1) were used. Figure 1 compares the turbulent kinetic energy (k) along the wall normal direction in the case of the RKE model with enhanced wall treatment technique. The results show grid independence as y<sup>+</sup> approached a value of one in the cells adjacent to the wall. The coarse grids produced peaks in k near the wall that appear closer to the true DNS solution. This occurs if the cell next to the wall is in the buffer region for the models with EWT. However, the trend is better-matched with the fine grids. Similar behavior was seen for the other high-Re models (RNG, SKE and RSM-LPS); hence grid independence plots for other models are not

presented. All models obtained grid independence with a  $139(\text{non-uniform}) \times 10 \times 10$  grid, so this grid was used for evaluation of these models. For the models using the standard and non-equilibrium wall function approaches, the first cell center next to the wall should be placed in the range of  $30 \le y^+ \le 500$  and, arbitrary grid refinement close to the wall is not appropriate. Hence, only uniform grids of  $30 \times 10 \times 10$  with  $y^+$  in cells next to the wall being in the range of 35-40 are used for models with these wall functions.

For low-Reynolds number flows (cases 2-5), the number of cells required to satisfy near-wall  $y^+>30$  is too small to be accurate. Hence, standard and non-equilibrium wall function approaches were not evaluated for low-Re flows. Only low-Re models (Abid, LB, LS, YS, AKN, and CHC) or high-Re models (like SKE, RNG (with low-Re differential viscosity model), RKE, and RSM-linear pressure-strain) with enhanced wall treatment are considered. Two uniform (50x10x10 and 80x10x10) and one non-uniform (100x10x10) grids were used for RKE, SKE, RNG, and RSM-LPS models with enhanced wall treatment to ascertain grid independency. The same grids were also used for the RSM-S $\omega$  (with low-Re correction) model. Figure 2 shows the turbulent kinetic energy for different grids predicted by SKE with enhanced wall treatment. Similar behavior was seen by other models as well. As the grid is refined to 100 non-uniformly-spaced cells, the results show very good grid independence. Hence this grid is used in all subsequent computations of low-Re cases with these models. For the square duct, the same grid is used in both the wall-normal directions (i.e. 100 x 100 x 10 cells).

Grid-convergence tests were also systematically done for each of the six low-Re k- $\epsilon$  models (Abid, LB, LS, YS, AKN, and CHC). Figure 3 shows one plot of turbulent kinetic energy in the Abid model for three different grids. All low-Re k- $\epsilon$  models were observed to achieve grid independence with 120 cells in the wall normal direction (giving a near-wall y<sup>+</sup> between 0.55-0.9). Hence this grid is used in all subsequent computations of low-Re cases with these models. In square duct flows, the same grid resolution of 120 cells is used in both wall-normal directions (i.e. 120x120x10).

# **4.3.** Computational Costs

Due to their varying complexities and convergence rates, both the total and per-iteration computational times for each model were different. Table 4 summarizes the time per iteration and total number of iterations to final convergence required by FLUENT (using 6 cores of a Dell Precision T7400 workstation with 2.66 GHz Intel Xeon processor and 8 GB RAM) with different models. As expected, the two equation models RKE, RNG and SKE with enhanced wall treatment require nearly the same time (per iteration as well as total time). On a per-iteration basis, the various two equations models are 5-30% less expensive than RSM-LPS (which solves 7 transport equations) with enhanced wall treatment. However, to obtain final converged results, RSM-LPS model is ~13-26 times more expensive. With standard and non-equilibrium wall functions, the two equation models are about 20-30% less expensive than RSM-LPS when compared on a per iteration basis but the time required to final convergence by RSM-LPS model reduces and it is only slightly more expensive. It seems that with finer grids, RSM-LPS model becomes increasingly expensive to achieve final convergence relative to two equation models. The enhanced wall treatment and standard/non-equilibrium wall functions are almost equally expensive for the same grid, but the grid required for enhanced wall treatment is finer. In all models tested, the computational requirement increases almost linearly with the grid size. Surprisingly, low-Re RSM-Sw model, which also solves 7 equations, is only about twice as expensive as the two equation models. All low-Re k- $\varepsilon$  models take nearly the same time per iteration, but the total times for LB and LS models are smaller. YS model took five times more time than LB and LS.

# **5. RESULTS AND DISCUSSION**

Results are first presented for non-MHD flows to show the accuracy of the various models without magnetic field. From these, models giving the best agreement are evaluated for the MHD flows after incorporating the changes due to the magnetic field effects.

## 5.1. High-Reynolds Number Non-MHD channel flow (Re=45818)

Figure 4 compares the turbulent kinetic energy predicted by the various models with the DNS data of Satake e al [48] for the grid independent 139x10x10 grid with enhanced wall treatment. It is seen that all models (RKE, RNG, SKE, and RSM-LPS) give nearly the same distribution of the turbulent kinetic energy. They underestimate the DNS peak values near the wall by 22-27%.

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Error decreases with distance from the wall, and TKE in the central core is predicted within 10%. Figure 5 shows similar behavior comparing models with standard wall functions. As theoretically required, the near-wall  $y^+$  has been maintained around 36-37. The results with standard wall functions were nearly the same as with the non-equilibrium wall functions probably because of the lack of flow separation or pressure gradient effects in a channel flow. As seen with the enhanced wall treatment, the peak TKE was again under-predicted, this time by even a larger amount (42%). The agreement in the core region is much better for all models, with the realizable model (RKE) giving slightly lower predictions.

The non-dimensionalized mean axial velocities predicted with the SKE and RSM-LPS models using enhanced wall treatment and standard wall functions are presented in Fig.6. The velocity profiles with non-equilibrium wall functions are not presented as they were nearly the same as with standard wall functions. It is seen that the enhanced wall treatment with  $y^+=1$  resolves velocity accurately all the way up to the viscous sublayer and matches best with the DNS results across the whole channel. Both models performed equally well with enhanced wall treatment, with errors consistently within 3%. With standard wall functions, as  $y^+$  is maintained ~36, the cell next to the wall stays in log-law region. Again both models predicted mean velocities well, although error with the RSM-LPS model increased to ~5% in the central core.

The Reynolds normal stresses predicted by the RSM-LPS model with all 3 wall treatments are compared with the DNS data in Figs. 7(a) and (b). With standard and non-equilibrium wall functions, the predictions matched closely with the DNS data in the core region except for the wall normal velocity fluctuations, which were underpredicted. The errors increased towards the wall especially in the axial and wall normal velocity fluctuations. Both wall functions performed equally but both missed the peak values close to the wall in all the three velocity fluctuations. The peak value of the RMS of axial velocity fluctuations is underpredicted by ~36% while the error in transverse and spanwise velocity fluctuations is smaller. The RMS of spanwise velocity fluctuations matched best with the DNS. The RSM-LPS model with enhanced wall treatment performed better than with standard or non-equilibrium wall functions in predicting all three velocity fluctuations, as expected. Again, the spanwise velocity fluctuations were predicted most accurately followed by wall normal fluctuations. The error in predicting peak value of axial

velocity fluctuations reduced from ~36% to ~12% by using the enhanced wall treatment. Overall, RSM-LPS with enhanced wall treatment predicted the anisotropy of Reynolds normal stresses reasonably well.

The mean streamwise pressure gradient predicted by various models is compared with the DNS data in Table 5. Overall, all models predicted the frictional losses within 10% error. The models with enhanced wall treatment (EWT) predicted the frictional losses more accurately than models using the standard (SWF) and non-equilibrium (NEWF) wall functions.

## 5.2. Low-Reynolds Number Non-MHD channel flow (Re=4586)

We next consider the low-Reynolds number non-MHD channel flow for which the various low-Re turbulence models are first evaluated. Figure 8 compares the turbulent kinetic energy predicted by various low-Re k- $\varepsilon$  models (Abid, LB, LS, YS, AKN, and CHC) with the DNS. The LS model greatly overpredicted throughout the domain, while the CHC model underpredicted near the wall and matched near the core. The 4 remaining models predicted similar values, matching the DNS data within 15% error near the wall but over-predicting (by ~60%) in the core. Overall, the LB model performed the best of all models, The YS model gave the correct trend across the whole domain, consistently overpredicting by 7-30%. The best low-Re k- $\varepsilon$  models (LB, AKN, and YS) are evaluated for mean axial velocity predictions in Figure 9. All three models predicted the mean axial velocity profile across the channel very well (within 5% error).

In addition to the low-Re k- $\varepsilon$  models, high Reynolds number k- $\varepsilon$  with EWT (RKE, RNG with differential viscosity, and SKE) and RSM models (RSM-LPS with EWT and RSM-S $\omega$  low-Re) also have been evaluated in this low-Re non-MHD channel flow. Figure 10 compares TKE predicted by these models. All models, except RNG and RSM-S $\omega$ , performed similarly by matching the peak values but over-predicting the values significantly (by ~120%) in the core. The RNG model overpredicted slightly more in the core than other models. RSM-S $\omega$  model matched TKE better in the core. Figure 11 compares the RMS of velocity fluctuations predicted by low-Re RSM-S $\omega$  and RSM-LPS model with the DNS. The RSM-S $\omega$  model, although it predicted the TKE best in the core, did not capture the anisotropy of Reynolds stresses even qualitatively. Because it was outperformed by the RSM-LPS model, the RSM-S $\omega$  model was not

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considered further in this study. The RSM-LPS model with enhanced wall treatment captured anisotropy qualitatively in all velocity fluctuations but overpredicted in the core. Figure 12 shows the comparison of the mean axial velocities given by RKE, SKE, and RSM-LPS models. All matched the DNS data closely except for some underprediction in the core.

Table 5 presents the mean streamwise pressure gradient predicted by various models. The best prediction of pressure gradient is by low-Re LB model (within 2% error) followed by the RSM-S $\omega$  model (3% error). All high Reynolds number k- $\varepsilon$  models with enhanced wall treatment overpredicted the pressure gradient by ~10%. The LS and CHC models were unreasonable, with frictional loss errors of ~95% and -16%. The other low-Re models predicted friction loss within 7%.

# 5.3. Low-Reynolds Number MHD channel flow (Re=4710, Ha=6)

The models (LB, SKE, and RSM-LPS) which performed better in low-Reynolds number non-MHD channel flow were then tested in low-Reynolds number MHD channel flow at a Reynolds number of 4710 and Ha = 6.0. Comparison of the computed TKE using the selected turbulence models with and without inclusion of the MHD sources/sinks (implemented through a userdefined function) is shown in Figure 13. The LB low-Re k- $\varepsilon$  model with MHD sources/sinks matches the DNS computed turbulent kinetic energy quite well in the core but underpredicts the high values close to the wall calculated by the DNS. The peak TKE is seen to be better predicted by LB without the MHD sources. The effect of the MHD sources/sinks on suppressing turbulence is clearly seen. SKE and RSM with enhanced wall treatment matched the peak values closely but overpredicted greatly (by 300-500%) in the core. Figure 14 presents the performance of RSM-LPS model in predicting anisotropy of Reynolds stresses and effect of magnetic field. Qualitatively, the distributions of the normal stresses are as in the DNS, however, the lateral stresses are over-predicted in the core and the axial velocity fluctuations are under-predicted close to the wall. The MHD source terms had little effect. It is interesting to note that although there is no MHD source/sink to the normal stress parallel to the magnetic field, there is an indirect effect through the other components. The models using enhanced wall treatment show very little effect of MHD source terms. This is likely due to the one equation model of Wolfstein [35] used in the viscous region in EWT which does not directly incorporate the MHD sources.

This contrasts with the strong effect observed in the low-Re LB model, where the source terms are applied throughout the domain.

Figure 15 compares the axial velocity in wall coordinates. The LB low-Re k- $\varepsilon$  model with MHD sources gives the best agreement with DNS data. However, part of profile in between 15<y<sup>+</sup><80 is under-predicted. This behavior in mean velocity is consistent with the behavior of the model in predicting lower TKE in around the same range of y<sup>+</sup>. The second best prediction is from the LB model without MHD sources. The predictions of RSM and SKE are similar, with the RSM-LPS performing slightly better. The underprediction of the normalized velocity in the core is mainly due to the higher frictional losses leading to higher friction velocity. The SKE and RSM models with enhanced wall treatment do not show much effect of MHD sources in the mean velocity. Figure 16 compares the axial velocity, as in Figure 15, but this time non-normalized mean velocity as a function of distance from the wall in the wall normal direction. The close match of predictions from all models with the DNS reinforces the assertion that the higher frictional losses are causing the differences in predictions in Figure-15.

We next examine the MHD source/sink terms in the k-equation and compare their magnitude with those extracted from the DNS budgets (Fig. 17). The trends predicted by all 3 models are reasonable, but the LB low-Re k- $\varepsilon$  model matches best with the DNS (within 20%). Although, the standard k- $\varepsilon$  model predicts the peak closely, it overpredicts the values in the core by ~300%. Interestingly, none of the models capture the small positive peak very close to the wall.

Figure 18 presents the sink term due to magnetic field in the turbulent dissipation rate ( $\epsilon$ ) equation. All 3 models correctly predict the asymptotic decay to zero dissipation in the core. The LB low-Re model correctly predicts the profile qualitatively across the whole channel but underestimates the values. The SKE and RSM models predict qualitatively similar profiles with negative peaks at y<sup>+</sup>~10. The SKE model gives the closest match although errors approach 50% near the wall.

Figures 19 and 20 give comparisons of the magnetic field source/sink terms in Reynolds normal stresses obtained by RSM-LPS. For S<sup>M+</sup><sub>ww</sub>, RSM behaves similar to the turbulent kinetic energy

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source. It underpredicts the peak value and overpredicts in the core. The positive values, which indicate a source in  $S^{M+}_{ww}$  below y<sup>+</sup><5, are again missed by the model. The MHD sink in  $S^{M+}_{uu}$  is qualitatively captured but the values are over-predicted across the whole length.

For this case, the LB model with MHD sources predicts the pressure gradient closest to the DNS (within 2.5%) followed by LB without MHD sources (Table 5). The SKE and RSM models overpredict the pressure gradient by about 20%. Adding the MHD sources improve the predictions slightly.

# 5.4. Low-Reynolds Number Non-MHD square duct flow (Re=5466)

The models are next evaluated for the fully-developed turbulent flow in a square duct bounded by four walls. For this case, it is well-known that the anisotropy in the Reynolds stresses generates cross-stream flows [16], which are not present in the laminar case. Turbulence models based on isotropic eddy-viscosity cannot predict such secondary flows [16]. To predict the secondary flows, it is necessary to use either non-linear/anisotropic two equation models [53-56], Reynolds-stress models [57], or algebraic stress models [58-59]. Hence, models other than the above are not expected to be accurate. However, they have been considered in this study to assess their inaccuracy and to evaluate their relative performance against the more expensive RSM. Figure 21 presents the comparison of turbulent kinetic energy along vertical bisector in a non-MHD square duct using LB, RKE, SKE and RSM-LPS models. The grid in all models resolved the flow up to the viscous sublayer  $(y^+ \sim 1)$ . The LB model predicts the turbulent kinetic energy better than other models. However, all models give excessive turbulent kinetic energy in the core region by over 100%. Figure 22 compares the predicted root mean square values of velocity fluctuations by the RSM model along vertical bisector of the duct with data from the DNS. RSM-LPS model with enhanced wall treatment, even when used with near wall spacing of  $y^+$ <1.1, over-predicts all the components of Reynolds normal stresses in the core by 40-75%. The agreement is however better in the near-wall region.

Figure 23 compares the mean axial velocity along the vertical bisector obtained by the different models. The realizable k- $\epsilon$  (RKE), standard k- $\epsilon$  (SKE) and LB models show similar reasonable behavior, as they agree with the DNS within ~8%. All 3 models overpredict in-between the wall

and the core and underpredict in the core region. The RSM model expectedly is slightly better but matches the other models in underpredicting the core region. Compared to the channel, the square duct flow is predicted with less accuracy, probably as a result of the inability to predict the secondary flows. Figure 24(a) and (b) show the mean axial velocity contours and secondary velocity vectors obtained by the DNS and the RSM-LPS model. Only the RSM model predicts the secondary flows, and hence results of other models are not shown. The bulging of the axial velocity profile is not predicted to the extent observed in the DNS.

Table 5 compares the mean streamwise pressure gradient predictions. It can be seen that the best model from the pressure gradient prediction is the LB model, which has an error of only  $\sim$ 7%. The RSM gives the highest pressure gradient ( $\sim$ 25% higher than DNS). SKE and RKE overpredict by  $\sim$ 12.0%.

# 5.5. Low-Reynolds Number MHD square duct flow (Re=5602, Ha=21.2)

The final test case considered is the MHD square duct flow, which is an appropriate geometry for the industrial application of electromagnetics to control flow in the continuous casting of steel. Here again, both isotropic viscosity-based models and the RSM models are evaluated, realizing still that the former cannot predict even qualitatively the cross-stream flow fields. Because of the magnetic field effects, for a square duct, the profiles of various quantities differ between the vertical and the horizontal bisectors. Hence profiles are compared along both these directions. Although the calculations of the channel flow were performed only using the magnetic induction method available in FLUENT, in the square duct flow, both magnetic induction and electric potential methods have been tested. The maximum magnitude of the induced magnetic field in the current simulations is only 0.039% of the externally applied field, so the magnetic induction method and electric potential method give virtually identical results.

Figures 25 and 26 compare the turbulent kinetic energies along vertical and horizontal bisectors respectively obtained from various models and the results of the DNS. It can be seen that MHD suppresses the turbulence energy more along the vertical bisector than along the horizontal bisector and only the LB model with MHD sources is able to predict this trend, reasonably matching with DNS (generally within 50%). The results with LB without MHD sources

overpredict the DNS data by 100-500%. The MHD sources/sinks proposed by Kenjereš and Hanjalić [12-13] provide significant improvements by predicting the correct trend of turbulence suppression, especially using the LB model. Both the RKE and RSM models over-predict the turbulence energy in the core along both the bisectors by ~500%. Moreover they do not capture the strong differential suppression of turbulence along the two bisectors, as was seen in the DNS and in the results of LB model with MHD sources. On the horizontal bisector close to the side walls, turbulence is not suppressed much because the induced current is parallel to the magnetic field in this region. The RKE and RSM models predict the peak value of the turbulent kinetic energy better along the horizontal bisector. Surprisingly, the RSM model is found to perform the worst among the tested models for suppressing turbulence by magnetic field effects.

Figure 27 and 28 respectively present the Reynolds stresses predicted by the RSM-LPS model and compare them with those of the DNS for the vertical and horizontal bisectors. It can be seen that the RSM model, as expected, can capture the anisotropy as well as the qualitative trends of Reynolds stresses although it overpredicts the values. The closest agreement is achieved along the horizontal bisector close to side walls where the effect of Lorentz force is the weakest.

Figure 29 and 30 give the mean axial velocity predictions from various models compared with the DNS along vertical and horizontal bisectors respectively. The DNS solution shows less flattening along the vertical bisector, which shows the importance of the secondary flows and the anisotropic suppression of turbulence by the magnetic field. All tested models predict about the same velocity profile along both bisectors. They match the DNS within ~4% along the horizontal bisector and overpredict velocity flattening along the vertical bisector by ~30%. The LB and RSM-LPS models are no better than the other models. MHD sources produce higher velocities, due to suppressing turbulence somewhat, but the agreement with DNS is not improved. The magnetic induction and the electric potential methods give the same mean axial velocity profiles along the two bisectors.

Figure 31 presents mean axial velocity contours and mean secondary velocity vectors in the cross-section. As shown by the DNS, the mean axial velocity contours and the secondary flows are significantly altered in the presence of the transverse magnetic field. The secondary velocities,

rather going into corners, now go towards the top and bottom walls, thus lifting the axial velocity contours in these regions towards the top and the bottom walls. After hitting the walls, these secondary flows move parallel to the top and bottom walls before turning towards the core at the center and thus cause a strong bulging in mean axial velocity there. This effect of strong bulging is not seen close to the side walls. It can be seen that none of the models is able to capture this effect. Although RSM predicts secondary flows, the differential effect of the magnetic field close to the top /bottom walls and the right/left side walls is missing. RSM predicts almost symmetric mean secondary and axial velocities except for a slight elongation of mean axial velocity (i.e. flattening) in the vertical direction. As mentioned earlier, LB and RKE do not predict secondary flows at all and over-predict the velocity flattening in the vertical direction, as also seen in the line plots of Fig. 29. Both the k- $\varepsilon$  models (LB and RKE) predict similar axial velocity across the cross-section.

Figure 32 and 33 show the MHD sources/sinks in the turbulence energy equation computed by the various models. The velocity-electric potential gradient correlation acts as a source whereas the Reynolds normal stresses perpendicular to the magnetic field act as sinks, as shown in the DNS data. The sink is stronger than the source giving a net effect of suppressing the turbulence. It can be seen that the LB model predicts this source reasonably correctly, followed by RKE and then RSM-LPS. The predictions are better along the stronger Lorentz force bisector. Both the RKE and the RSM-LPS over-predict the MHD sources to TKE along both bisectors.

Because of the MHD effects, there is no longer bisector symmetry as in the non-MHD square duct. Consequently, the friction factors along the bottom-horizontal and left-vertical walls are different, as shown in Fig. 34. Along the bottom horizontal wall, the friction factor shows two side peaks with a large dip at the center. Along left-vertical wall, the friction factor shows a central flat region with two side dips. None of the models is seen to predict these trends correctly. Both the k- $\epsilon$  models (LB and RKE) give similar profiles, with a central overpredicted peak. The RSM-LPS model predicts the side peaks with a central dip along both walls but does not completely agree with the DNS results. RSM suggests larger frictional losses, especially in the corners. The best agreement is seen with LB model with MHD sources. The LB model, without MHD sources, overpredicts friction along both walls.

Finally, a comparison of mean streamwise pressure gradients computed by the various models again shows (Table 5) that the LB low-Re model with MHD sources performs best by matching within ~2% error with the DNS predictions. LB model without MHD sources is next, followed by RKE with EWT. The performance of these models is similar in the non-MHD and the MHD cases.

#### 6. CONCLUSIONS

In this study several turbulence models of k- $\varepsilon$  and Reynolds stress transport category are evaluated for their ability to predict turbulent flow fields subjected to a magnetic field. Five test cases of flows in a channel and square duct have been computed and the results are compared with DNS data. The MHD sources/sinks in k- and  $\varepsilon$ - equations for k- $\varepsilon$  models and in Reynolds stresses for RSM, as proposed by Kenjereš and Hanjalić [12-13], were implemented through UDFs in the FLUENT code. The performance of these models, on the basis of their predictions of mean velocities, RMS of velocity fluctuations, turbulent kinetic energy, MHD sources and frictional losses can be summarized as follows:

In both high and low Reynolds number channel flows, all of the models predicted mean axial velocity reasonably well (within 5% error), given fine-enough grids for grid-independence (EWT and low-Re) or satisfaction of the y+ requirements (SWF and NEWF). However, the turbulent kinetic energy was much less accurate, often exceeding 60% overprediction in the core. In high Reynolds number channel flows, models underpredicted near-wall peak turbulence energy whereas in low-Reynolds number channel flows, they showed better agreement near the wall but over-predicted values in the core. For the MHD flows, the implementation of the MHD sources improved predictions for low-Re k- $\varepsilon$  models. The high-Re models which use the wall treatments did not show much improvement with MHD sources, perhaps due to the lack of MHD effects in the wall formulations.

In the case of low-Re square duct flows, the models tested did not predict the mean axial velocities to a good accuracy (error ranging ~8-30%) because of the secondary flows generated due to turbulence anisotropy. The turbulent kinetic energy was overpredicted in the core, often

exceeding ~60%, by all models except LB in MHD duct. The effect of turbulence suppression by magnetic field was not properly captured on mean velocity, Reynolds stresses/turbulent kinetic energy and frictional losses by any single model in a MHD duct, even after inclusion of the MHD sources of turbulence.

For problems involving high Reynolds number, the SKE model offers reasonable accuracy at low computational cost. Adding enhanced wall treatment improves accuracy slightly over standard wall laws, but significantly increases cost. For flows with low Re number, the Lam-Bremhorst (LB) low-Re k- $\epsilon$  model performed better than the others in both hydrodynamic and magnetic field influenced turbulent flows. Given the need to compute complex industrial flows with efficient computational use, using these 2 models with appropriate changes for magnetic field effects provides a reasonable compromise of accuracy and speed. Finally, the RSM-LPS model with enhanced wall treatment offers similar accuracy with the added ability of capturing turbulence anisotropy and secondary flows, but its computational cost is very high.

# ACKNOWLEDGMENTS

The authors gratefully thank the Continuous Casting Consortium, Department of Mechanical Science & Engineering, University of Illinois at Urbana-Champaign, IL, and the National Science Foundation (Grant CMMI-07-27620) for providing support for this work. The CU-FLOW code was developed in the Computational Fluid Dynamics Laboratory, UIUC with the support of a NASA GSRP fellowship to Aaron Shinn. We would like to thank Aaron Shinn for providing the basic CU-FLOW code and the DNS data in non-MHD square duct flow. The FLUENT code was provided by ANSYS, Inc.

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Model	$f_1$	$f_2$	$f_{\mu}$	$\mathcal{E}_w/k_w$ (wall BC)
Abid	1.0	$\left(1 - \frac{2}{9}\exp(-\operatorname{Re}_{T}^{2})/36\right)^{2}$	$\tanh\left(0.008\mathrm{Re}_{y}\right)\left(1+4\mathrm{Re}_{T}^{-3/4}\right)$	$\varepsilon_w = v \frac{\partial^2 k}{\partial y^2} / k_w = 0$
		$\left(1-\exp\left(-\operatorname{Re}_{y}/12\right)\right)$		
LB	$1 + (0.05 / f_{\mu})^3$	$\left(1-\exp(-\operatorname{Re}_{T}^{2})\right)$	$(1 - \exp(-0.0165 \operatorname{Re}_y))^2 (1 + 20.5 / \operatorname{Re}_T)$	$\frac{\partial \varepsilon}{\partial y} = 0 / k_w = 0$
LS	1.0	$\left(1-0.3\exp(-\operatorname{Re}_{T}^{2})\right)$	$\exp(-3.4/(1 + \text{Re}_T/50)^2)$	$\varepsilon_w = 0 / k_w = 0$
YS	$\frac{\sqrt{\mathrm{Re}_T}}{1+\sqrt{\mathrm{Re}_T}}$	$\frac{\sqrt{\mathrm{Re}_T}}{1+\sqrt{\mathrm{Re}_T}}$	$ \begin{pmatrix} (1+1/\sqrt{Re_T}) \\ (1-\exp\left(-1.5x10^{-04} \operatorname{Re}_y \\ -5x10^{-07} \operatorname{Re}_y^3 - 10^{-10} \operatorname{Re}_y^5) \end{pmatrix} \right)^{1/2} $	$\varepsilon_{w} = 2\nu \left(\frac{\partial\sqrt{k}}{\partial y}\right)^{2} / k_{w} = 0$
AKN	1.0	$ \left(1 - 0.3 \exp(-\left(\operatorname{Re}_{\tau}/6.5\right)^{2}\right) \right) $ $ \left(1 - \exp(-\operatorname{Re}_{\varepsilon}/3.1)\right)^{2} $	$(1+5.0/\operatorname{Re}_{T}^{3/4} \exp(-(\operatorname{Re}_{T}/200)^{2}))$ $(1-\exp(-\operatorname{Re}_{\varepsilon}/14))^{2}$	$\varepsilon_w = v \frac{\partial^2 k}{\partial y^2} / k_w = 0$
CHC	1.0	$ \begin{pmatrix} 1 - 0.01 \exp(-\operatorname{Re}_{T}^{2}) \\ (1 - \exp(-0.0631 \operatorname{Re}_{y})) \end{pmatrix} $	$(1 - \exp(-0.0215 \operatorname{Re}_y))^2$ $(1 + 31.66 / \operatorname{Re}_T^{5/4})$	$\varepsilon_w = v \frac{\partial^2 k}{\partial y^2} / k_w = 0$
Where	e, $\operatorname{Re}_T = \frac{\rho k^2}{\mu \epsilon}$	$\frac{2}{\epsilon}$ , $\operatorname{Re}_{y} = \frac{\rho \sqrt{k} y}{\mu}$ and $\operatorname{Re}_{\varepsilon} = \frac{\rho}{\mu}$	$\frac{(\mu \epsilon / \rho)^{1/4} y}{\mu}$	

Table 1 Damping functions and wall boundary conditions for different low-Re k- $\epsilon$  models

Table 2 Various ter	ms and constant	of lov	v-Re k-e	models

Model	D	Ε	$C_1$	<i>C</i> <sub>2</sub>	$\sigma_k$	$\sigma_{_{arepsilon}}$	$C_{\mu}$
Abid	0	0	1.45	1.83	1.0	1.4	0.09
LB	0	0	1.44	1.92	1.0	1.3	0.09
LS	$2\nu \left(\frac{\partial \sqrt{k}}{\partial y}\right)^2$	$2vv_t \left(\frac{\partial^2 \overline{u}}{\partial y^2}\right)^2$	1.44	1.92	1.0	1.3	0.09
YS	0	$VV_t \left(\frac{\partial^2 \overline{u}}{\partial y^2}\right)^2$	1.44	1.92	1.0	1.3	0.09
AKN	0	0	1.5	1.9	1.4	1.4	0.09
CHC	0	0	1.44	1.92	1.0	1.3	0.09

Geometry	Re	Grid	Comput.	Spatial resolution	На	
,		$(N_x x N_y x N_z)$	Domain	$(\Delta x^+, \Delta y^+, \Delta z^+)$		$\mathbf{w}_{b}$ / $a p$ / $az$
		, _	(XxYxZ)			
Channel	45818	1024x1024x768	πx1x2.5π	9.16, 0.163-4.25, 17.2	0	20.45 / 2.0
(Case-1)	$(Re_{\tau}=1120)$					
	(Satake et al)					
Channel	4586	128x97x128	πx1x2.5π	7.36, 0.08-4.91, 18.4	0	15.28 / 2.0
(Case-2)	$(Re_{\tau}=150)$					
	(Iwamoto et al)					
Channel	4710	64x128x64	0.5πx1x1.25π	7.36, 0.08-4.9, 9.2	6.0	15.7 / 2.0
(Case-3)	$(Re_{\tau}=150)$					
	(Noguchi et al)					
Square	5466	160x160x1024	1x1x8	1.47-3.24, 1.47-3.24, 2.81	0	15.187 / 4.0
duct	$(Re_{\tau}=360)$			(1% stretching in x- and y-)		
(Case-4)	(Shinn et al)					
Square	5602	128x128x512	1x1x16	1.41-4.92, 1.41-4.92, 11.28	21.2	1.057/0.01857
duct	$(Re_{\tau}=361)$			(2% stretching in x- and y-)		
(Case-5)	(Chaudhary et al)					

Table 3 Various p	parameters in	different DNS	calculations	considered	during	evaluation	[48-52	2]
					<u> </u>	-	A	

Where, 
$$\operatorname{Re}_{\tau} = \frac{D_{1}u_{\tau}}{v}$$
,  $\operatorname{Re} = \frac{D_{2}W_{b}}{v}$ , and  $Ha = B_{y}D_{1}\sqrt{\frac{\sigma}{\rho v}}$ 

Channel:  $D_1 = \delta$ ,  $D_2 = 2\delta$  ( $\delta = 0.5$  is half channel height)

Square duct:  $D_1 = D_2 = D$ , (D = 1 is the side of the square duct)

Table 4 Comparison of the **time taken per iteration** (sec) / **# of iterations in final convergence** by FLUENT (parallel with 6 cores) with various models, wall treatment methods and Reynolds

numbers in non-MHD	channel flow	for	final	grids
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Turbulence Model	Wall treatment	Re=4586 Re=4586		Re=45818			
	method		100x10x10	139x10x10	30x10x10		
	En wall treatment	-	0.19/2289	0.22/3818	-		
RKE	Non-eq wall fn	-	-	-	0.11 / 1227		
	Std wall fn	-	-	-	0.11 / 1227		
	En wall treatment	-	0.19/ 2289	0.23 / 3195	-		
SKE	Non-eq wall fn	-	-	-	0.11/954		
	Std wall fn	-	-	-	0.11/954		
	En wall treatment	-	0.20 / 2700	0.24 / 3125	-		
RNG	Non-eq wall fn	-	-	-	0.11/954		
	Std wall fn	-	-	-	0.11/954		
RSM-LPS	En wall treatment	-	0.21 / 55033	0.29 / 38689	-		
	Non-eq wall fn	-	-	-	0.14 / 2464		
	Std wall fn	-	-	-	0.13/3115		
RSM-Sw	Low-Re RSM model	-	0.22 / 4568	-	-		
Abid	Low-Re k-ɛ model	0.20 / 5400	-	-	-		
LB	Low-Re k-ɛ model	0.20/3150	-	-	-		
LS	Low-Re k-ɛ model	0.20 / 3075	-				
YS	Low-Re k-ɛ model	0.20 / 17700	-	-	-		
AKN	Low-Re k-ɛ model	0.21/4571	-	-	-		
CHC	Low-Re k-ɛ model	0.21 / 5214	_	-	-		
CHC LOW-Re K-E model 0.21/3214							

# Table 5 Mean streamwise pressure gradient in different flows predicted by various models

			-			-
Channel	Channel	Channel		Square duct	Square	duct
(Re=45818)	(Re=4586)	(Re=471	0, Ha=6)	(Re=5466)	(Re=5602, Ha=21.2)	
		Mag-Inc	d Method		Mag-Ind/Elec F	ot Methods
		With MHD	Without		With MHD	Without
		sources	MHD		sources	MHD
			sources			sources
2.0	2.0	2	.0	4.0	0.018	57
1.97	2.20	-	-	4.46	0.0228/0.0228	-
2.02	2.20	2.40	2.47	4.46	-	-
1.99	2.20	-	-	-	-	-
2.08	2.16	2.37	2.42	5.0	0.0244	-
1.83	-	-	-	-	-	-
1.90	-	-	-	-	-	-
1.83	-	-	-	-	-	-
1.84	-	-	-	-	-	-
1.85		-	-	-	-	-
1.94		-	-	-	-	-
1.89		-	-	-	-	-
1.85	-	-	-	-	-	-
-	1.94	-	-	-	-	-
-	2.07	-	-	-	-	-
-	1.97	2.04	2.18	4.28	0.0190	0.0215
-	3.87	-	-	-	-	-
-	2.13	-	-	-	-	-
-	2.11	-	-	-	-	-
-	1.68	-	-	-	-	-
	Channel (Re=45818) 2.0 1.97 2.02 1.99 2.08 1.83 1.90 1.83 1.84 1.85 1.94 1.97 1.97 2.02 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.99 2.08 1.83 1.83 1.90 1.97 1.97 1.99 2.08 1.83 1.84 1.85 1.94 1.85 1.94 1.85 1.94 1.85 1.94 1.85 1.85 1.94 1.85 1.85 1.94 1.85 1.94 1.85 1.94 1.85 1.85 1.94 1.95 1.94 1.95 1.95 1.95 1.95 1.95 1.95 1.95 1.95	Channel (Re=45818)Channel (Re=4586) $(Re=45818)$ $(Re=4586)$ $2.0$ $2.0$ $1.97$ $2.20$ $2.02$ $2.20$ $2.02$ $2.20$ $2.08$ $2.16$ $1.83$ - $1.90$ - $1.83$ - $1.84$ - $1.85$ - $1.94$ - $1.85$ - $1.94$ - $1.89$ - $1.89$ - $1.94$ - $ 1.94$ $ 2.07$ $ 3.87$ $ 2.13$ $ 2.11$ $ 1.68$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $

EWT: Enhanced Wall Treatment

SWF: Standard Wall Function

NEWF: Now-Equilibrium Wall Function

LPS: Linear Pressure Strain

So: Stress-Omega



# **Figure Captions**

Fig-1 Grid independence study in high Reynolds number channel flow for RKE with enhanced wall treatment

Fig-2 Grid independence study in low Reynolds number channel flow for SKE with enhanced wall treatment

Fig-3 Grid independence study in low Reynolds number channel flow for Abid low-Re k- $\epsilon$  model

Fig-4 Comparison of turbulent kinetic energy in various models with enhanced wall treatment in high Reynolds number channel flow

Fig-5 Comparison of turbulent kinetic energy in various models with standard wall function approach in high Reynolds number channel flow

Fig-6 Comparison of normalized mean axial velocity in SKE and RSM-LPS with standard wall functions and enhanced wall treatment in high Reynolds number channel

Fig-7 Comparison of RMS of velocity fluctuations in RSM-linear-pressure-strain with (a) nonequilibrium and standard wall functions (b) enhanced wall treatment in high Reynolds number channel flow

Fig-8 Comparison of turbulent kinetic energy predicted by low-Re k-ε models with the DNS in low Reynolds number channel flow

Fig-9 Comparison of the mean axial velocity predicted by low-Re k-ε models with the DNS in low Reynolds number channel flow

Fig-10 Comparison of turbulent kinetic energy predicted by RKE, RNG, SKE and RSM-LPS with enhanced wall treatment and low-Re RSM-S $\omega$  turbulence models with the DNS in the low Reynolds number channel flow

Fig-11 Comparison of RMS of velocity fluctuations by RSM models with the DNS in low Reynolds number channel flow

Fig-12 Comparison of mean axial velocity by SKE, RKE, RSM-LPS models with enhanced wall treatment with the DNS in low Reynolds number channel flow

Fig-13 Comparison of turbulent kinetic energy in low Reynolds number MHD channel flow with various models

Fig-14 Comparison of RMS of velocity fluctuations predicted by RSM-LPS with enhanced wall treatment with the DNS in low Reynolds number MHD channel flow

 Fig-15 Comparison of normalized mean axial velocity vs. normalized wall distance in wall units in low Reynolds number MHD channel flow in various models

Fig-16 Comparison of mean axial velocity vs. distance from the wall in low Reynolds number MHD channel flow in LB and SKE models

Fig-17 Comparison of the MHD source/sink in the k-equation / budget (DNS) in low Reynolds number MHD channel flow in various models with the DNS

Fig-18 Comparison of MHD sink in  $\varepsilon$ -equation / budget (DNS) in low Reynolds number MHD channel flow in various models with the DNS

Fig-19 Comparison of the MHD source/sink in  $\overline{w'w'}$ -equation / budget (DNS) in low Reynolds number MHD channel flow in RSM-LPS model with the DNS

Fig-20 Comparison of MHD source/sink in  $\overline{u'u'}$ -equation / budget (DNS) in low Reynolds number MHD channel flow in RSM-LPS model with DNS

Fig-21 Comparison of TKE predicted by various models with the DNS along vertical bisector in a non-MHD square duct

Fig-22 Comparison of RMS of velocity fluctuations predicted by RSM-LPS model with the DNS in non-MHD square duct along vertical bisector

Fig-23 Comparison of mean axial velocity predicted by various models with the DNS in non-MHD square duct along vertical bisector

Fig-24 Comparison of mean axial velocity contours and secondary velocity vectors in non-MHD square duct (a) DNS (Re=5466, Ha=0, Shinn et al [51]) (160x160x1024)

(b) RSM-linear-pr-strain, En wall treatment (Re=5466, Ha=0, 100x100x10)

Fig-25 Comparison of TKE in various models with the DNS in MHD square duct along vertical bisector

Fig-26 Comparison of TKE in various models with the DNS in MHD square duct along horizontal bisector

Fig-27 Comparison of RMS of velocity fluctuations predicted by RSM with the DNS in MHD square duct along vertical bisector

Fig-28 Comparison of RMS of velocity fluctuations predicted by RSM with the DNS in MHD square duct along horizontal bisector

Fig-29 Comparison of the mean axial velocity predicted by various models with the DNS in MHD square duct along vertical bisector

Fig-30 Comparison of mean axial velocity in various models with DNS in MHD square duct along horizontal bisector

Fig-31 Comparison of mean axial velocity contours and secondary velocity vectors in MHD duct

(a) DNS (Chaudhary et al [52]) (Re=5602, Ha=21.2, 128x128x512)

- (b) RSM, En wall treatment, Mag-Induction (Re=5602, Ha=21.2, 100x100x10)
- (c) Realizable k-ɛ, En wall treatment, Mag-Induction (Re=5602, Ha=21.2, 100x100x10)
- (d) LB, Low-Re k-ɛ, Mag-Induction (Re=5602, Ha=21.2, 120x120x10)

Fig-32 Comparison of MHD source/sink in k-equation / budget (DNS) predicted by various models with the DNS in MHD square duct along vertical bisector

Fig-33 Comparison of MHD source/sink in k-equation / budget (DNS) in various models with the DNS in MHD square duct along horizontal bisector

Fig-34 Comparison of the friction factor in MHD square duct along bottom-horizontal and left-vertical walls in various models with the DNS





Fig-1 Grid independence study in high Reynolds number channel flow for RKE with enhanced wall treatment



Fig-2 Grid independence study in low Reynolds number channel flow for SKE with enhanced wall treatment



Fig-3 Grid independence study in low Reynolds number channel flow for Abid low-Re k-ε model



Fig-4 Comparison of turbulent kinetic energy in various models with enhanced wall treatment in high Reynolds number channel flow



Fig-5 Comparison of turbulent kinetic energy in various models with standard wall function approach in high Reynolds number channel flow



Fig-6 Comparison of normalized mean axial velocity in SKE and RSM-LPS with standard wall functions and enhanced wall treatment in high Reynolds number channel



Fig-7 Comparison of RMS of velocity fluctuations in RSM-linear-pressure-strain with (a) non-equilibrium and standard wall functions (b) enhanced wall treatment in high Reynolds number channel flow



Fig-8 Comparison of turbulent kinetic energy predicted by low-Re k- $\epsilon$  models with the DNS in low Reynolds number channel flow



Fig-9 Comparison of the mean axial velocity predicted by low-Re k-ε models with the DNS in low Reynolds number channel flow



Fig-10 Comparison of turbulent kinetic energy predicted by RKE, RNG, SKE and RSM-LPS with enhanced wall treatment and low-Re RSM-Sω turbulence models with the DNS in the low Reynolds number channel flow



Fig-11 Comparison of RMS of velocity fluctuations by RSM models with the DNS in low Reynolds number channel flow

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Fig-12 Comparison of mean axial velocity by SKE, RKE, RSM-LPS models with enhanced wall treatment with the DNS in low Reynolds number channel flow



Fig-13 Comparison of turbulent kinetic energy in low Reynolds number MHD channel flow with various models



Fig-14 Comparison of RMS of velocity fluctuations predicted by RSM-LPS with enhanced wall treatment with the DNS in low Reynolds number MHD channel flow



Fig-15 Comparison of normalized mean axial velocity vs. normalized wall distance in wall units in low Reynolds number MHD channel flow in various models





Fig-16 Comparison of mean axial velocity vs. distance from the wall in low Reynolds number MHD channel flow in LB and SKE models



Fig-17 Comparison of the MHD source/sink in the k-equation / budget (DNS) in low Reynolds number MHD channel flow in various models with the DNS



Fig-18 Comparison of MHD sink in  $\epsilon$ -equation / budget (DNS) in low Reynolds number MHD channel flow in various models with the DNS



Fig-19 Comparison of the MHD source/sink in w'w' -equation / budget (DNS) in low Reynolds number MHD channel flow in RSM-LPS model with the DNS



Fig-20 Comparison of MHD source/sink in  $\overline{u'u'}$ -equation / budget (DNS) in low Reynolds number MHD channel flow in RSM-LPS model with DNS



Fig-21 Comparison of TKE predicted by various models with the DNS along vertical bisector in a non-MHD square duct



Fig-22 Comparison of RMS of velocity fluctuations predicted by RSM-LPS model with the DNS in non-MHD square duct along vertical bisector



Fig-23 Comparison of mean axial velocity predicted by various models with the DNS in non-MHD square duct along vertical bisector



Fig-25 Comparison of TKE in various models with the DNS in MHD square duct along vertical bisector



Fig-26 Comparison of TKE in various models with the DNS in MHD square duct along horizontal bisector



Fig-27 Comparison of RMS of velocity fluctuations predicted by RSM with the DNS in MHD square duct along vertical bisector



Fig-28 Comparison of RMS of velocity fluctuations predicted by RSM with the DNS in MHD square duct along horizontal bisector



Fig-29 Comparison of the mean axial velocity predicted by various models with the DNS in MHD square duct along vertical bisector



Fig-30 Comparison of mean axial velocity in various models with DNS in MHD square duct

along horizontal bisector



MHD duct



Fig-32 Comparison of MHD source/sink in k-equation / budget (DNS) predicted by various models with the DNS in MHD square duct along vertical bisector



Fig-33 Comparison of MHD source/sink in k-equation / budget (DNS) in various models with the DNS in MHD square duct along horizontal bisector



Fig-34 Comparison of the friction factor in MHD square duct along bottom-horizontal and

left-vertical walls in various models with the DNS